

Graz kHz SLR Station

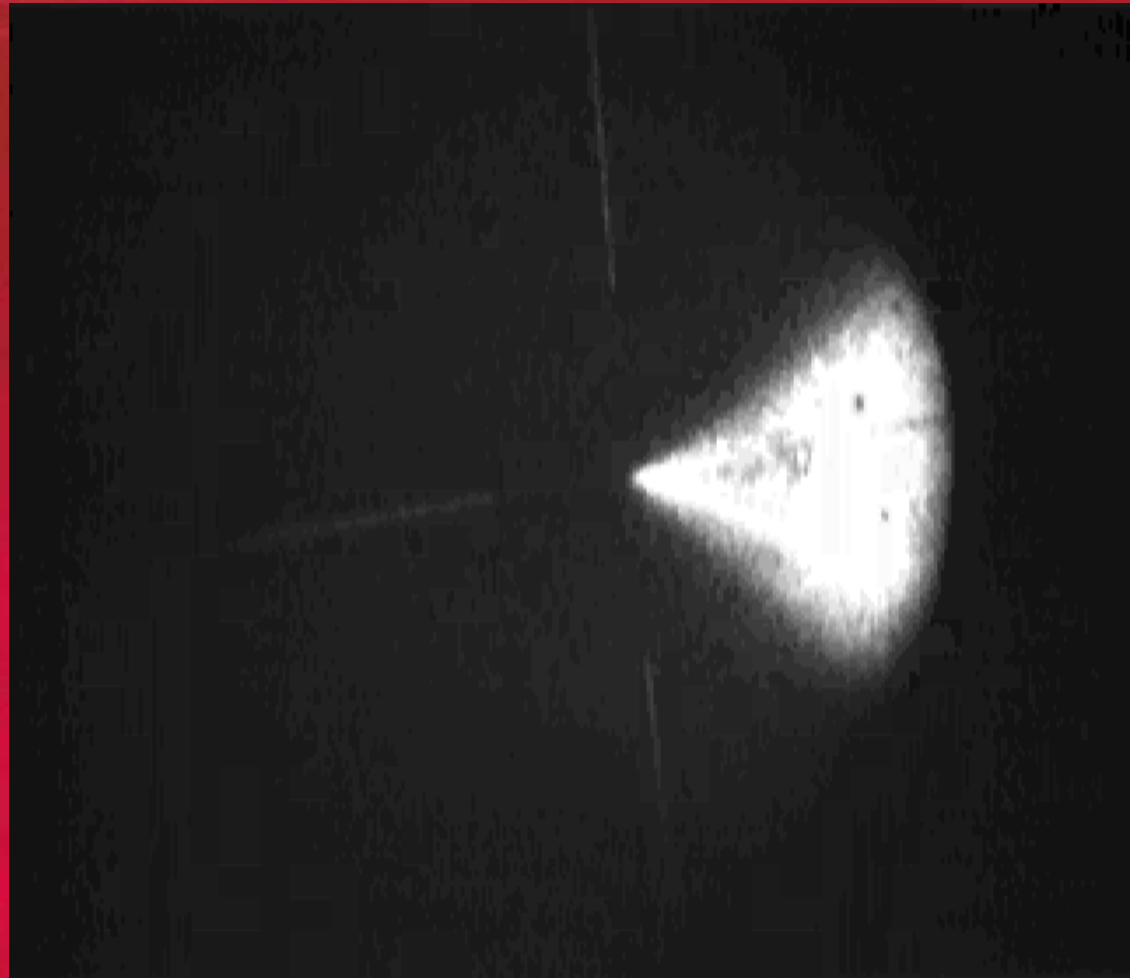
Measuring Atmospheric Seeing with kHz SLR

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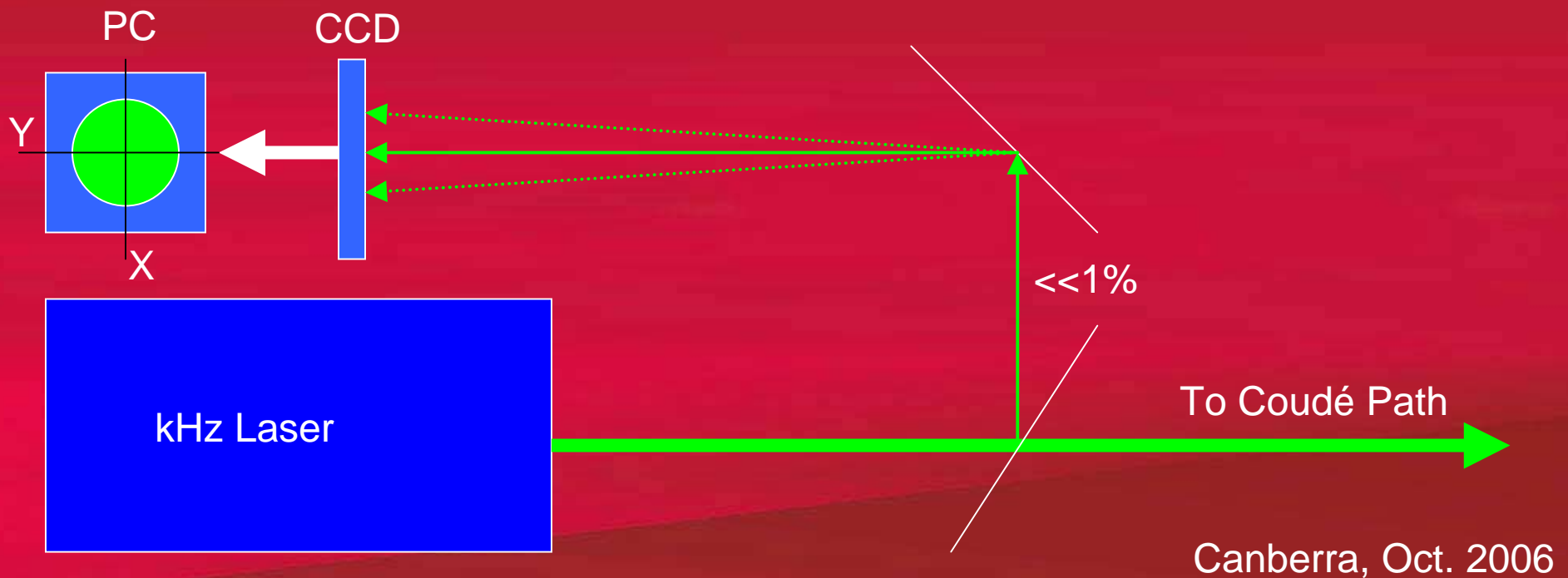
Beam Pointing: Not always stable ...

- Laser Beam, night time
- Backscatter, seen by ISIT
- 25 Frames per second
- Pointing is NOT stable
- Pointing Jitter: up to $50 \mu\text{rad}$
- Frequency of this wobbling:
Few Hz up to few 10 Hz



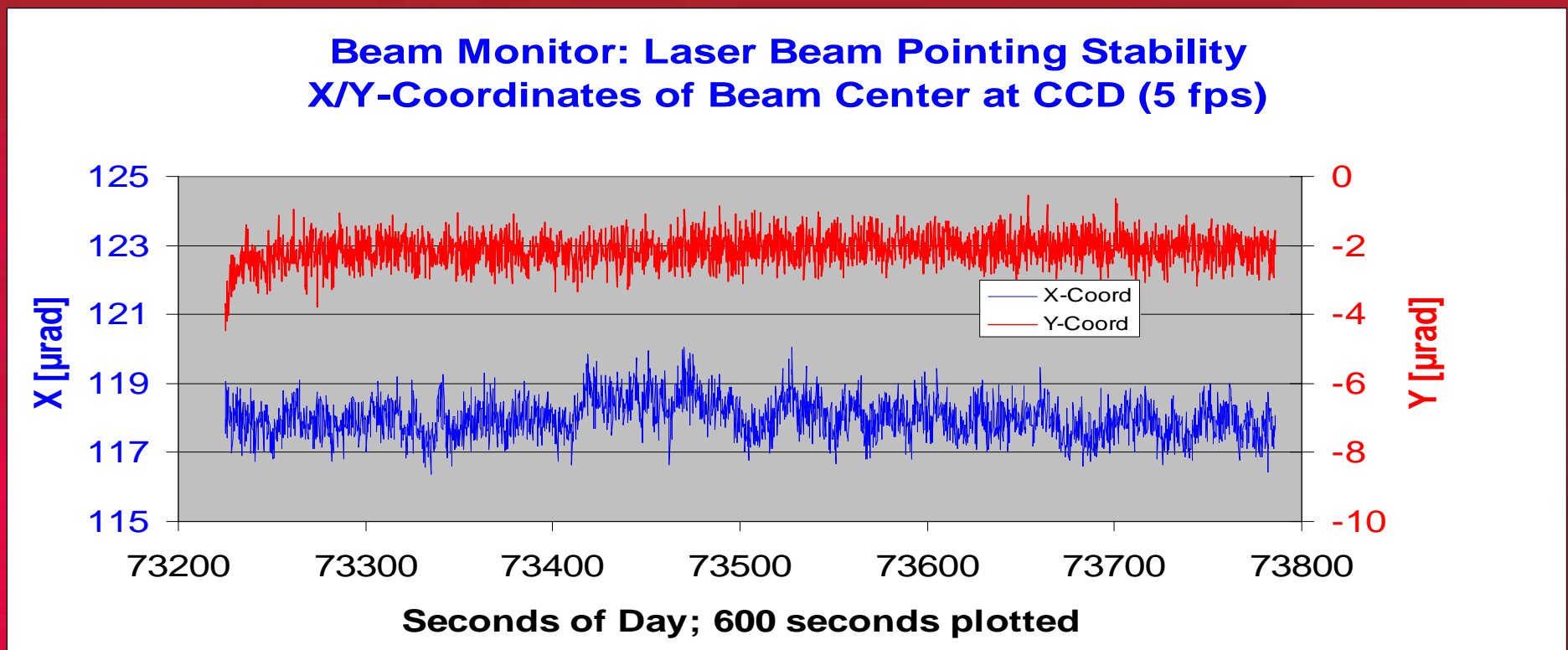
Possible Reason: Laser Beam Pointing ?

- We installed a **Laser Beam Monitor** at the laser box beam exit:
 - A mirror reflects a small portion ($\ll 1\%$) of the laser on a CCD chip;
 - CCD Image: Monitored by a PC, with up to 30 fps;
 - Center Coordinates of Beam (X / Y) calculated, and stored.



Possible Reason: Laser Beam Pointing ?

- Only few μrad ($\ll 1''$) wobble (mainly measurement accuracy);
- **Conclusion: Wobble is NOT caused by** ~~Even this DECREASED by following beam expander;~~
- **Laser Beam Pointing Instability** (not shown here)



Other possible reason: Atmosphere ...

- Beam Wobble caused by atmospheric micro-turbulences
=> Atmospheric „Seeing“:
- Expected Amplitudes: Some Arcseconds
 - Laser Beam Wobble: Up to $50 \mu\text{rad}$ (= $10''$)
 - Expected Frequencies: From few Hz up to few 10 Hz
 - Can be more than 100 Hz, but at decreasing amplitudes ...
 - Laser Beam Wobble: Up to 10 or 15 Hz visible at ISIT images

Seeing Effects for SLR ???

- Graz Beam Divergence ($< 10''$) and Seeing ($\approx 2'' - 7''$):
 - Both with Similar Magnitudes ...
 - Seeing for SLR can be WORSE than Astronomical Seeing:
=> Fast moving telescope, faster changing atmospheric conditions
- Degrades laser pointing accuracy
- May reduce return rates;
- No problem for LEOs, but may have effects for GPS etc.



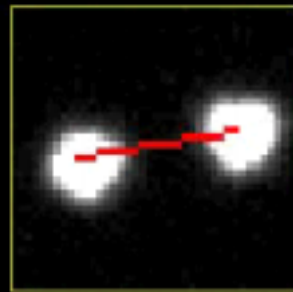
What worsens „Seeing“ ?

- Seeing Values are influenced by:
 - Actual atmosphere: Wind; layers, gradients etc.
 - Elevation of Satellite: Seeing is worse at lower elevation
 - Temperature differences: Seeing is worse in winter time:
 - Graz Observatory rooms are heated; but isolation is almost ZERO
 - This causes lot of turbulences around station in WINTER time ...
- Motion of telescope: Seeing is worse at higher speeds !!!

Measuring Seeing: Standard Method

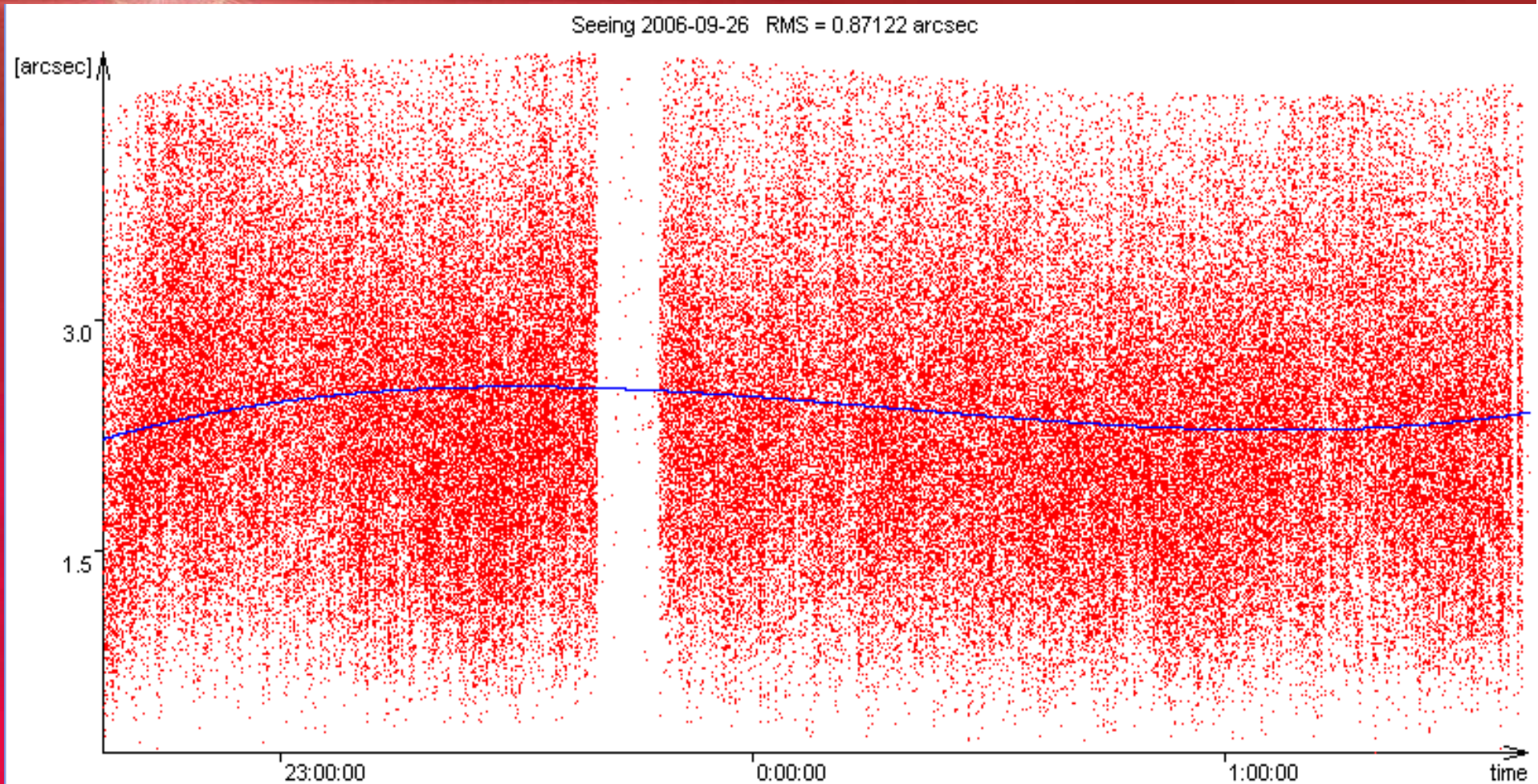
- Hartmann – Shack: 2 Holes at the entrance pupil;
- Observing Polar Star: Gives 2 spots on the CCD sensor;
- Variation of spot distances gives Seeing Value

Seeing = 3.57 arcsec



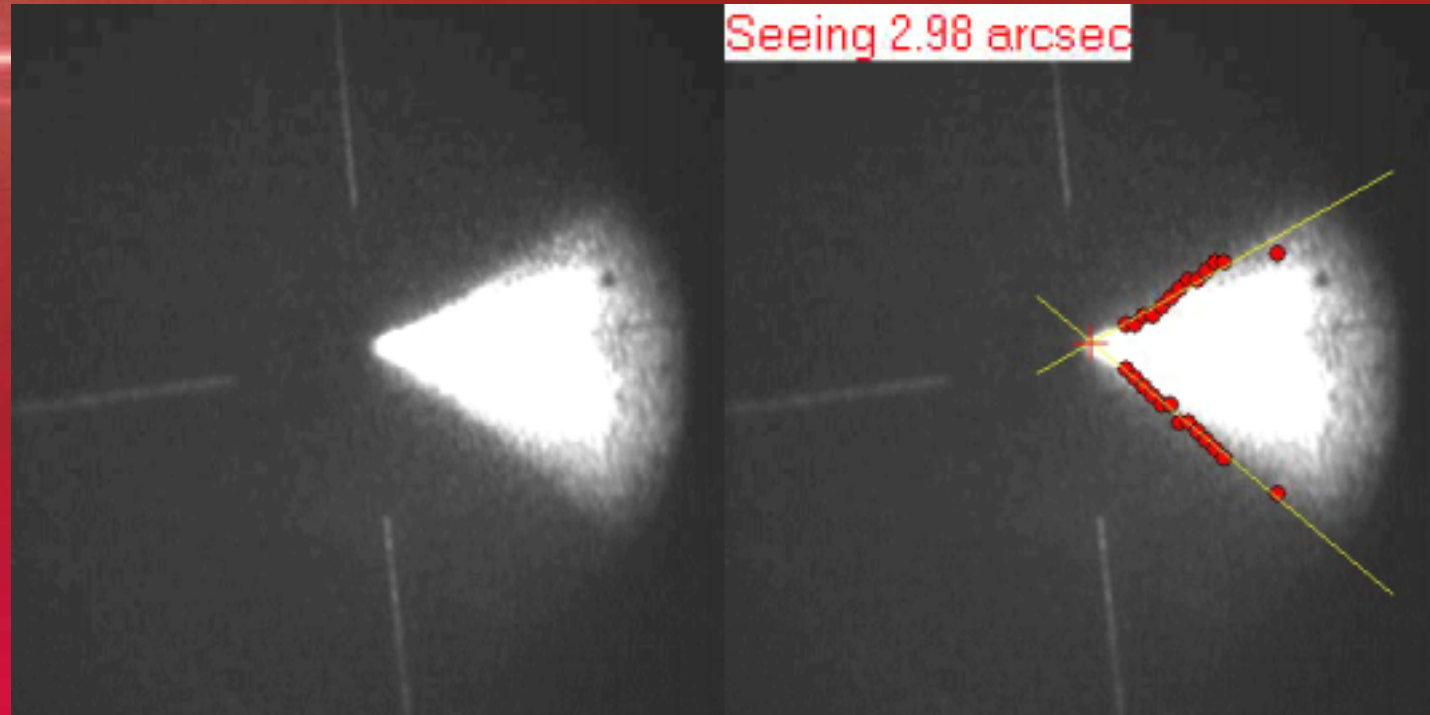
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Seeing Values with Hartmann-Shack



- Our „Best“ Seeing ever measured: Below 3“ during 3 hours

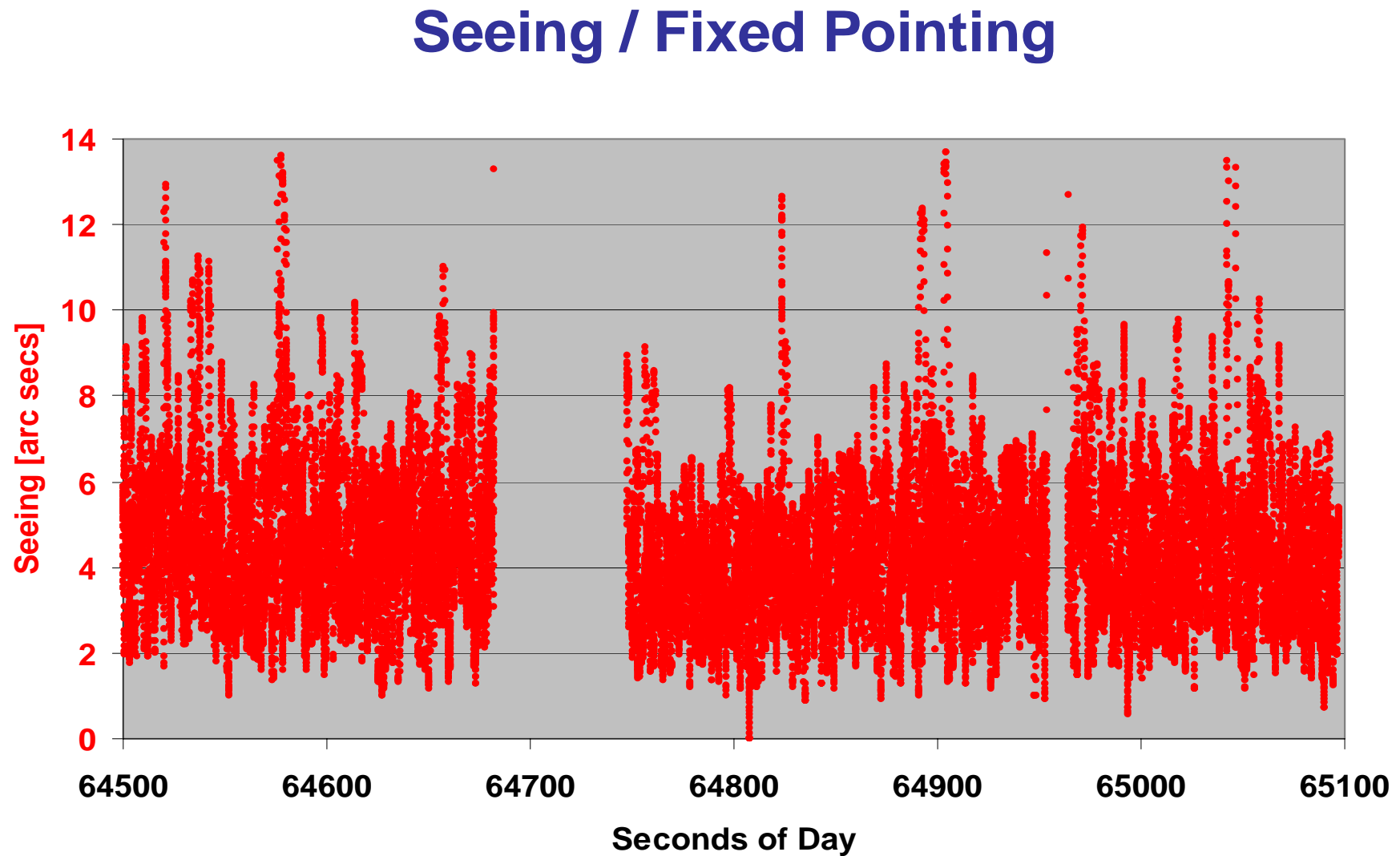
Deriving Seeing Values from kHz



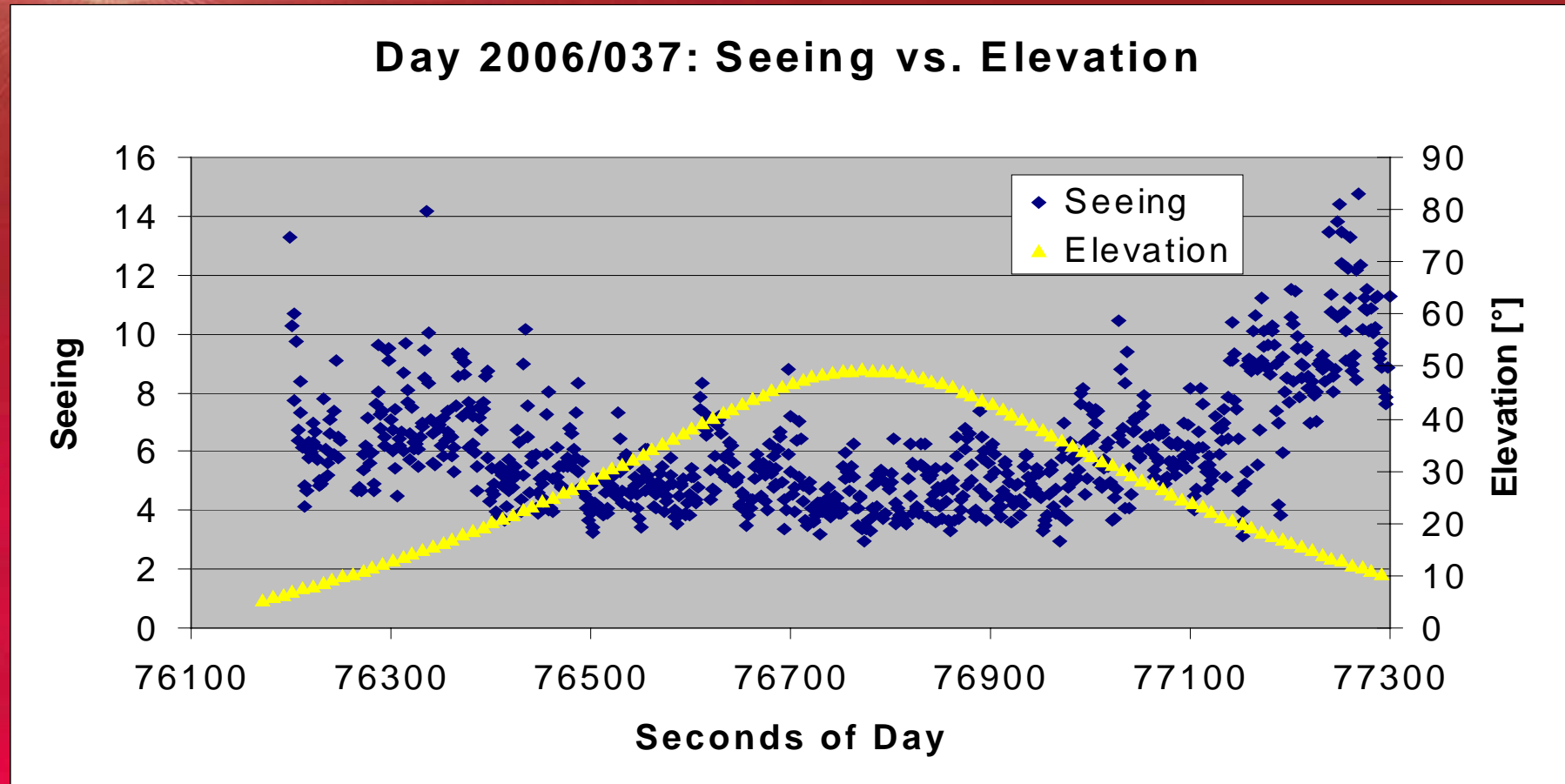
- Laser Beam Backscatter is monitored by ISIT; frames into PC
- Real Time Image Processing: Determine Peak of Laser Beam;
- Coordinates of Peak determine Seeing Area;
- FWHM of this area => Astr. Seeing (arc secs);
- Compare with Hartmann – Shack Results

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Seeing Values derived from kHz Laser



Seeing Values Derived from kHz Laser

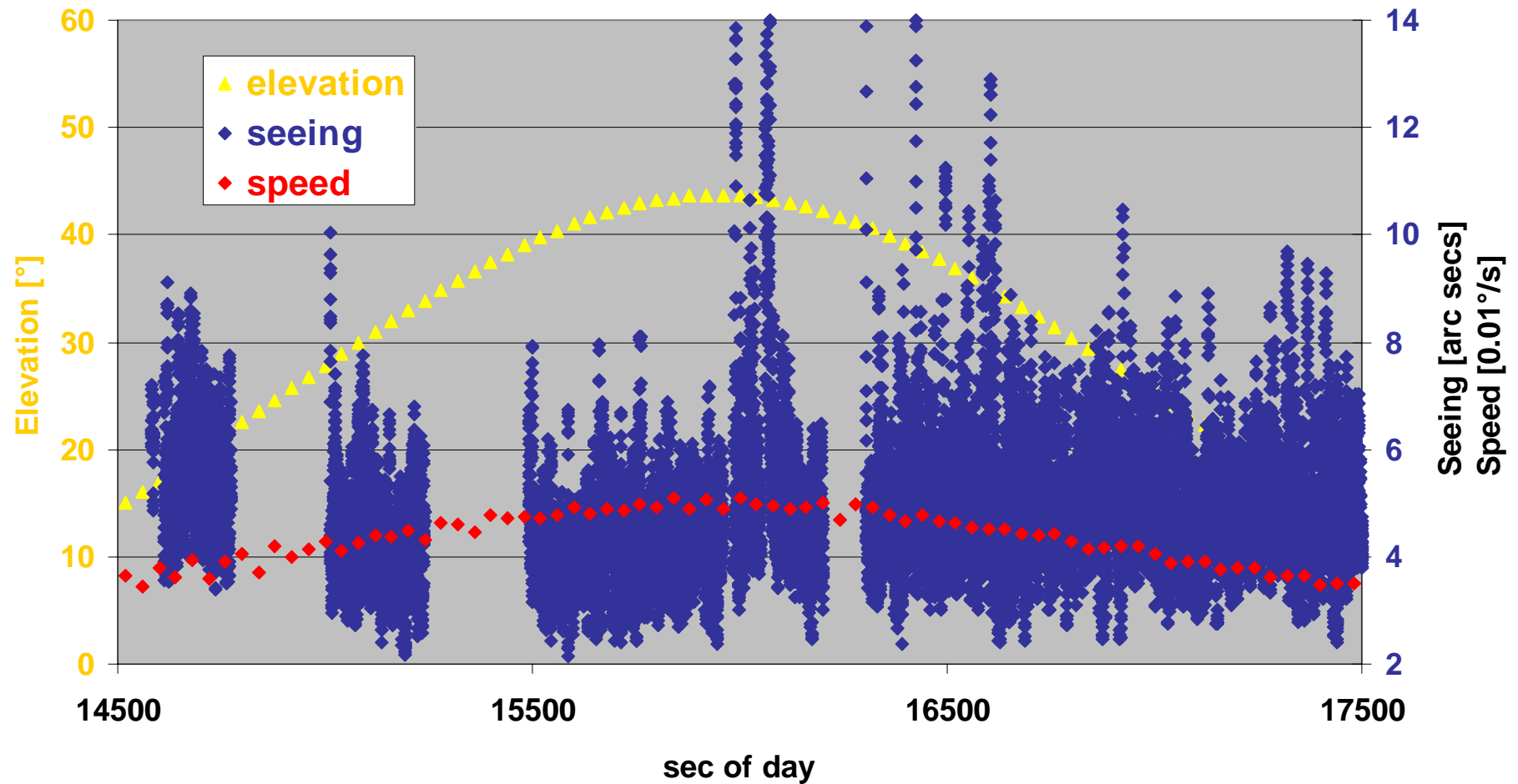


Ajisai / Day 037/2006: Seeing changes with elevation

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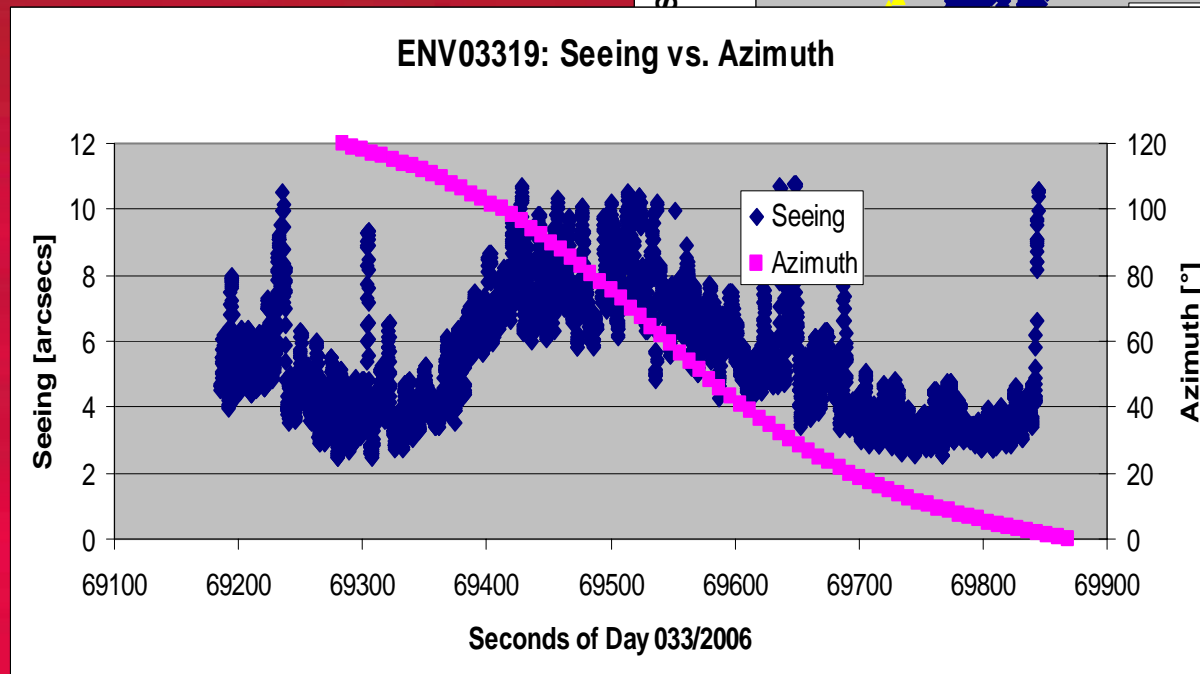
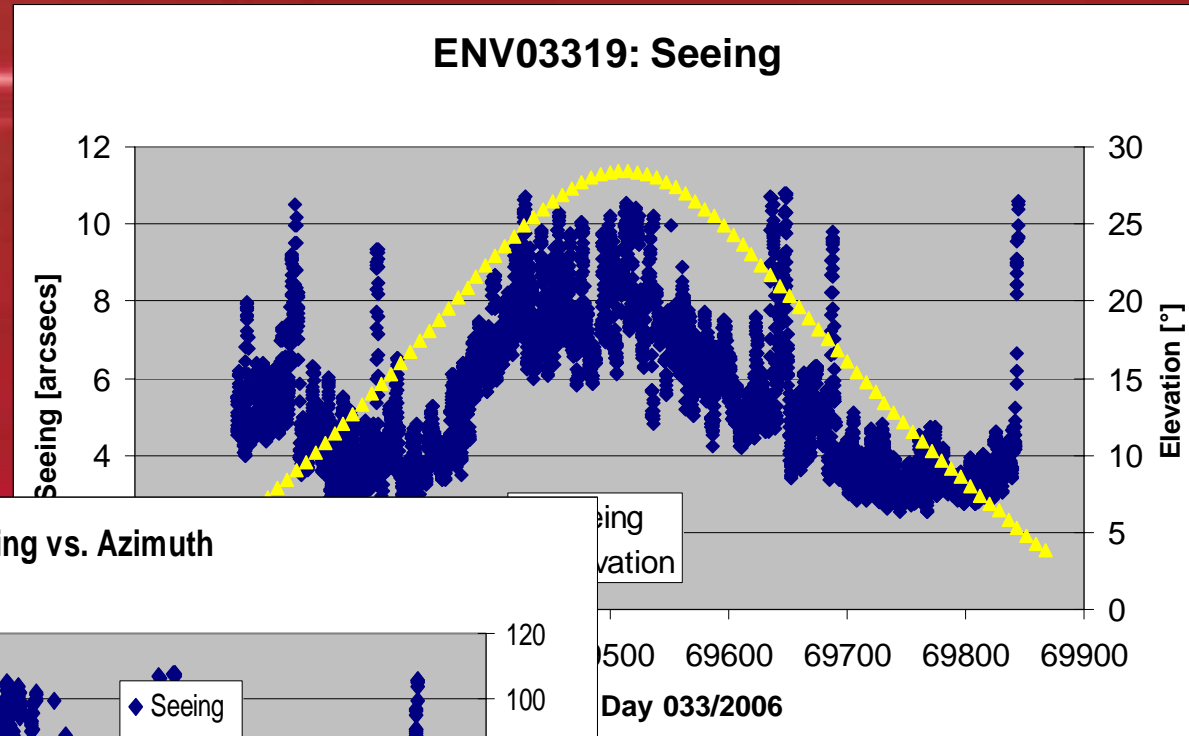
Seeing Values Derived from kHz Laser

Lageos1 322 03: 43° EL



Seeing Values Derived from kHz Laser

- Envisat: Day 033
- $< 30^\circ$ Elevation



- At 90° Azimuth:
=> Obs. Roof !!!
- Heating Influence

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Conclusions:

- It is possible to derive Seeing Values from a kHz Laser Beam;
- Seeing Influence is BIG enough to spoil Laser Beam Pointing;
- This might reduce Return Rate from High Orbiting Sats;
- Biggest Influence in Graz: Heating of Observatory
 - We have to live with it – or freeze to death ☺

- Plans to reduce effect with a Fast Steering Mirror, controlled by Seeing Offsets derived from kHz Beam (10" max, 40 Hz max);



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