

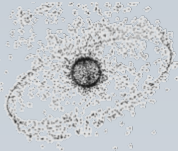
Space Debris Laser Ranging Activities in Zimmerwald

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ILRS Virtual Workshop 2023





Space Debris Laser in Zimmerwald

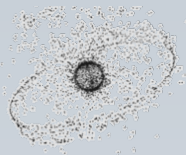
Design Considerations

- 47W, 200Hz laser system
- piggy-back on 0.8m ZimMAIN
 - no interference with geodetic SLR operations
 - requires separate transmit optics and new receiver chain
 - might be too heavy
 - requires new tracking and control software
- piggy-back on ZIMLAT
 - not an option because new 1 kHz geodetic laser system to be mounted piggy-back on ZIMLAT
- ZIMLAT Coudé
 - could profit from existing receiver chain and software
 - transmit/receive switch can handle 100/200Hz
 - time sharing with geodetic SLR operations

→ preferred solution to start with

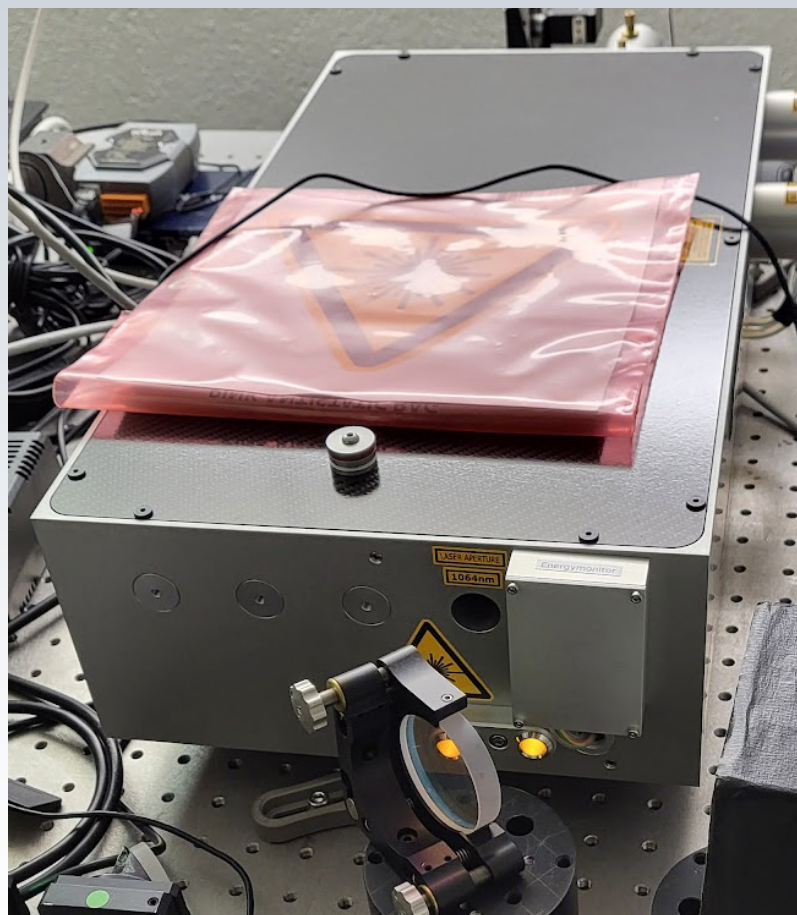
(bistatic system to be evaluated in future)





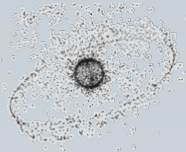
Space Debris Laser System

INNOLAS EVO II-200 Laser specifications

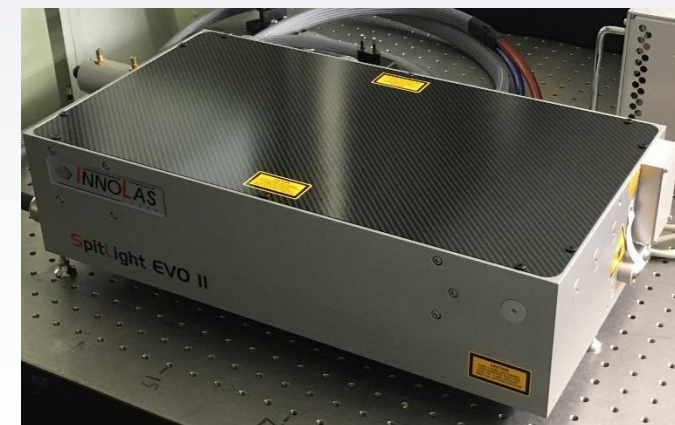
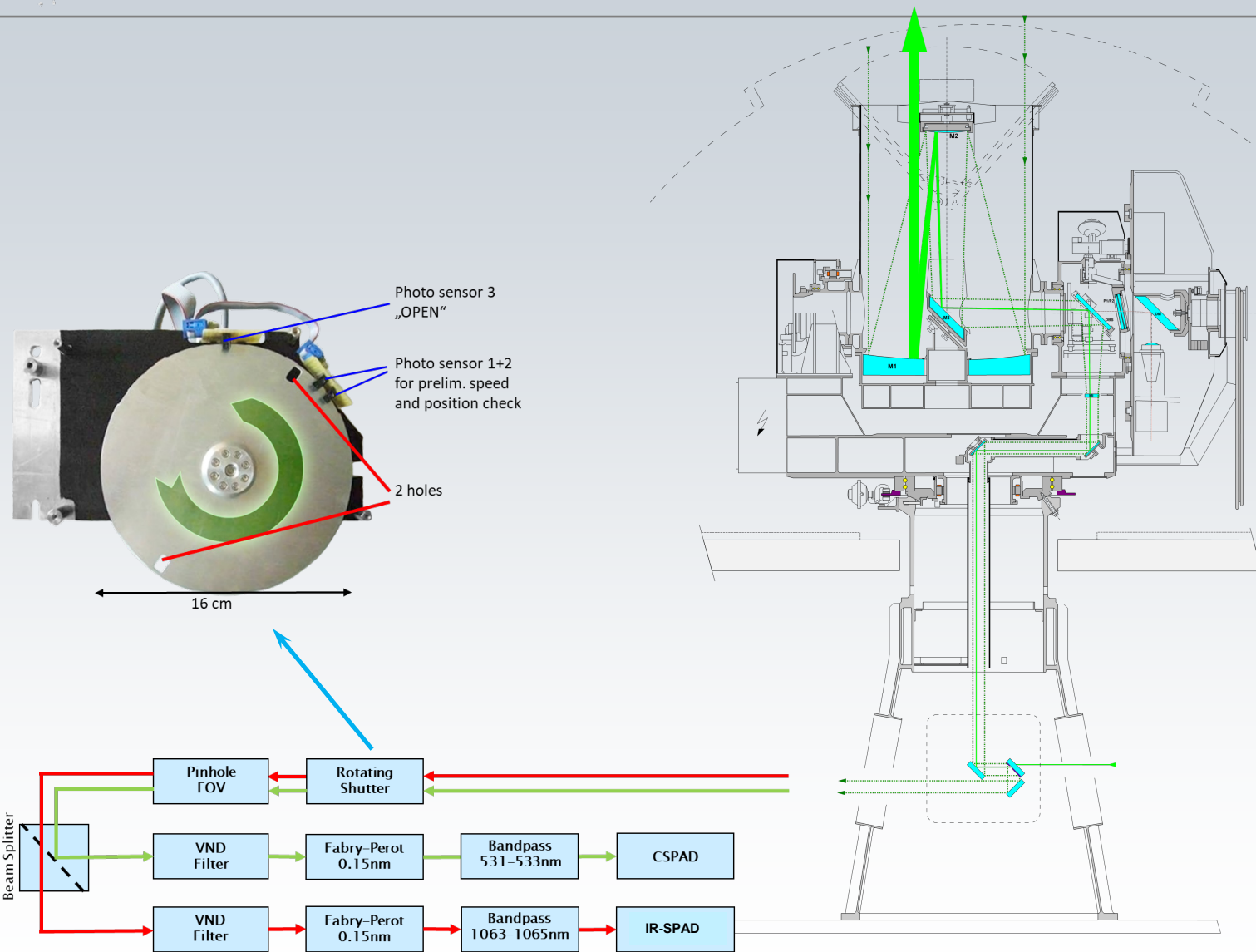


Wavelength	1064 nm
Pulse repetition rate	200 Hz (100 Hz subsampling)
Pulse length (min)	7 ns
Pulse energy (adjustable)	<1 mJ - 236 mJ
Pulse-to pulse timing jitter	± 1 ns
Average power (@200Hz)	< 0.2 W - 47 W
Beam diameter (1/e)	8mm
Beam divergence	<0.5 mrad
Polarization	Linear (horizontal)

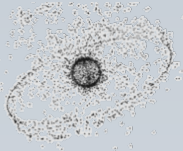
ZIMLAT Coudé Configuration



T. Schildknecht et al.: Space Debris Laser Ranging Activities in Zimmerwald
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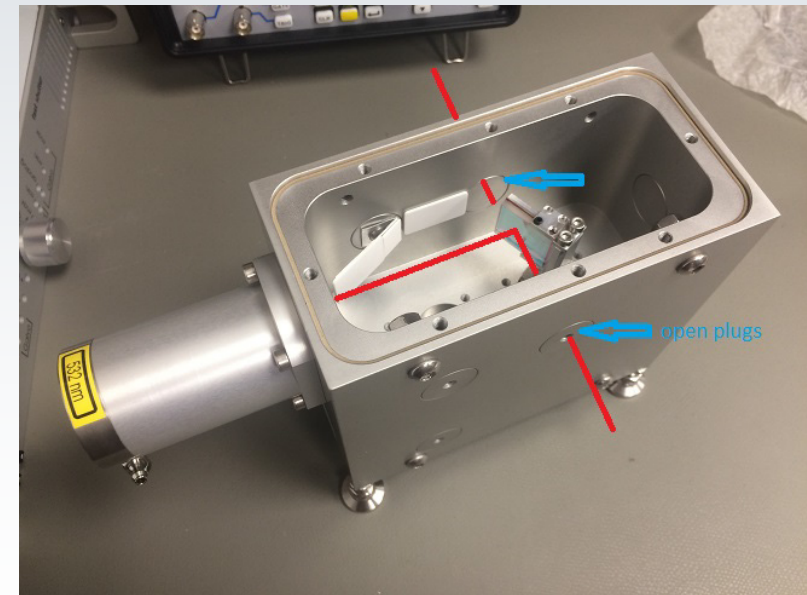
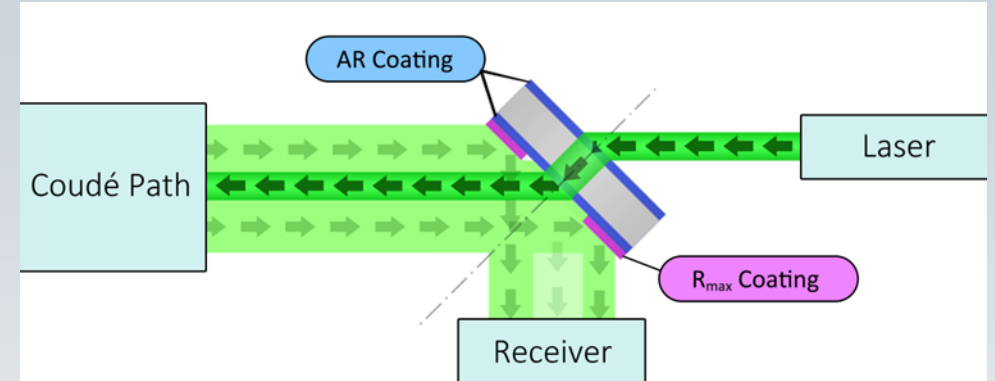
Challenges of a Monostatic Coudé Configuration

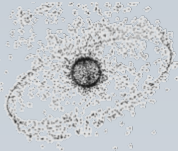
Calibration Measurements

- transmit/receive switch open during calibration
→ stray light/reflections from the telescope enter receiver
- **need to attenuate transmission beam**
(also required for eye safety when using terrestrial target)

Beam Attenuation

- $< 10\mu\text{J}$ required
- **$4 \cdot 10^{-5}$ attenuation**
 - 10^{-2} internal attenuation
 - 10^{-3} external attenuation

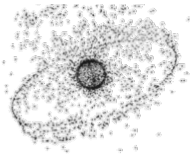




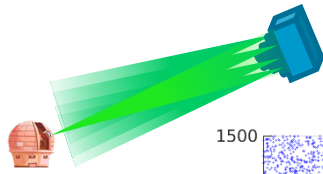
Challenges of a Monostatic Coudé Configuration

Transmit – Receive Conflicts

- problem: time of flight $\sim x$ * pulse repetition interval
→ laser transmitted while receive switch open (would destroy receiver)
- standard procedure:
 - change laser repetition rate from 100Hz to 109Hz (200Hz to 218Hz)
(no issue for geodetic 1W laser)
- requires 6 separate working points
 - laser parameters needed to be optimized for each working point
(thermal conditions!)
 - 6 working points:
 - 101.5 Hz (± 1.5 Hz)
 - 104.5 Hz (± 1.5 Hz)
 - 107.5 Hz (± 1.5 Hz)
 - 193.2 Hz (± 3.4 Hz)
 - 200.0 Hz (± 3.4 Hz)
 - 206.8 Hz (± 3.4 Hz)
- “afterglow” of lenses in the transmit/receive path not yet quantified

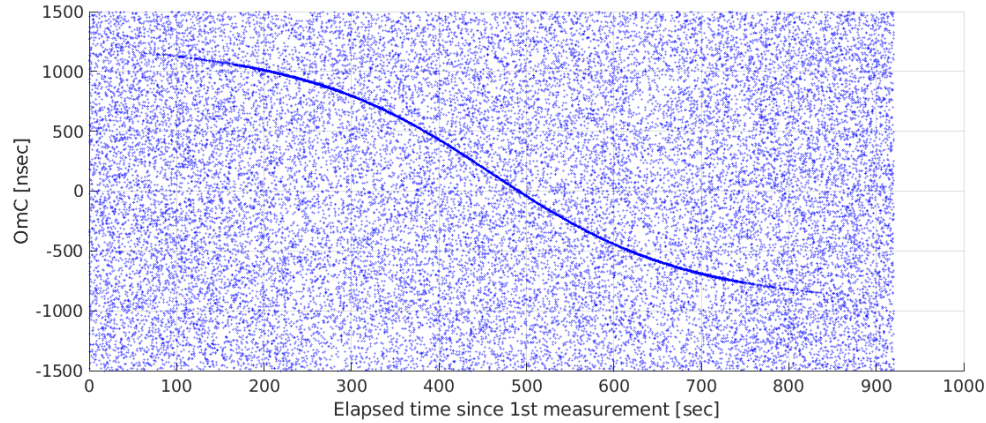


Processing Weak Signals from SDLR



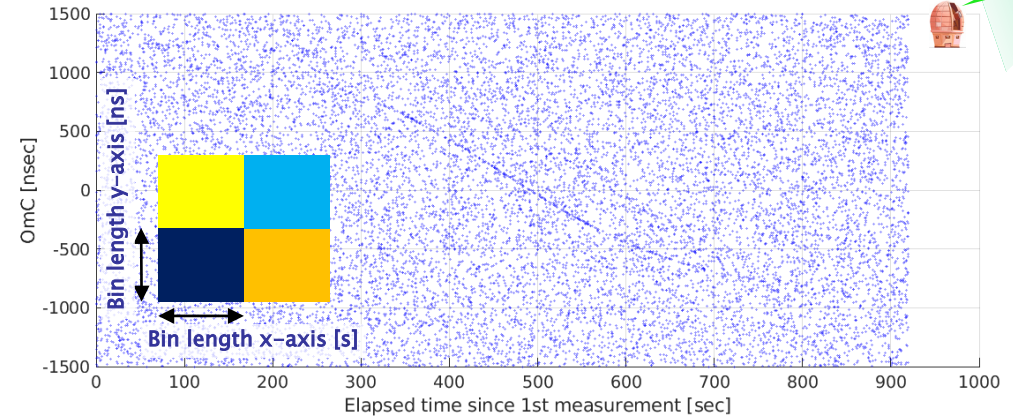
Case A: High Signal-to-Noise

100Hz

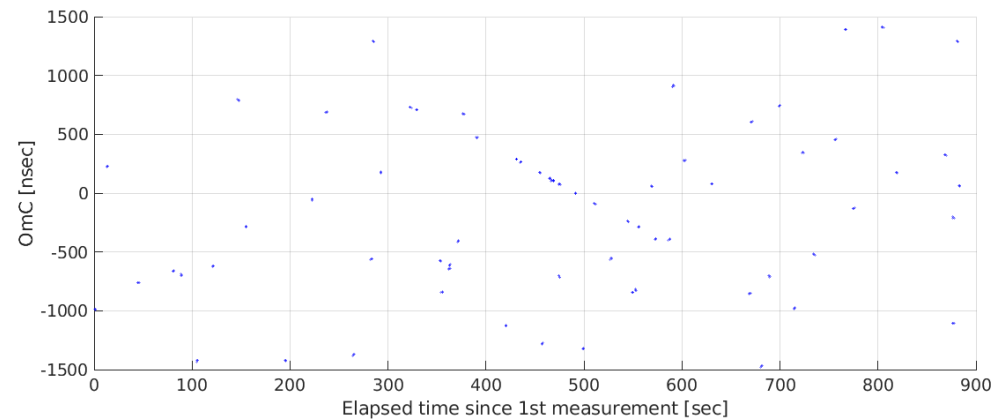
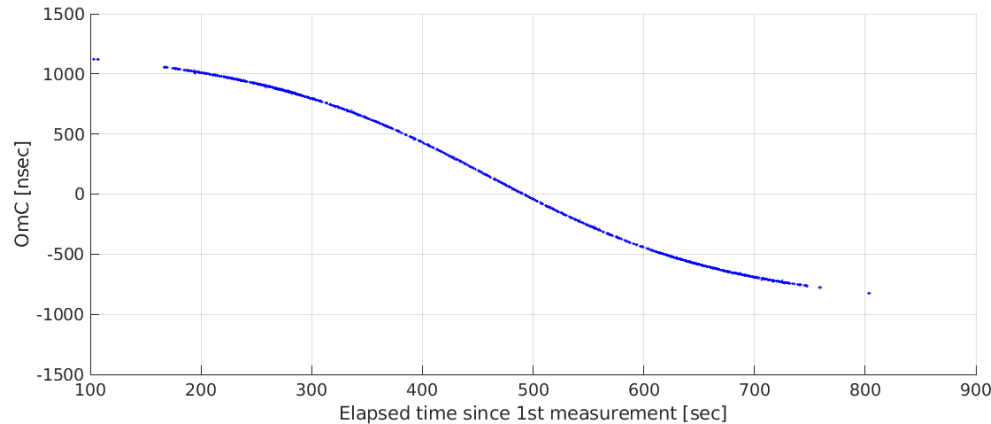


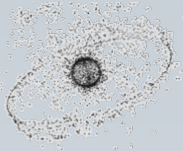
Case B: Low Signal-to-Noise

100Hz



Filtered signal with traditional methods

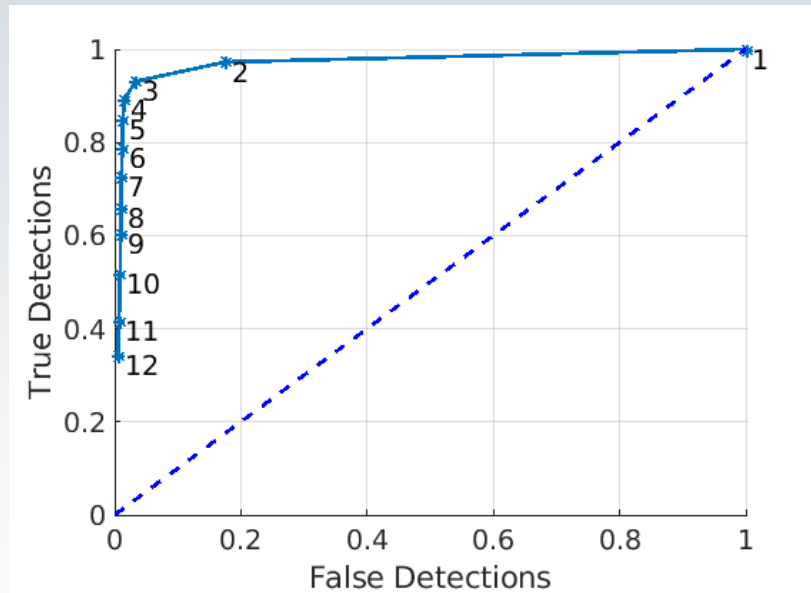




Processing Weak Signals from SDLR

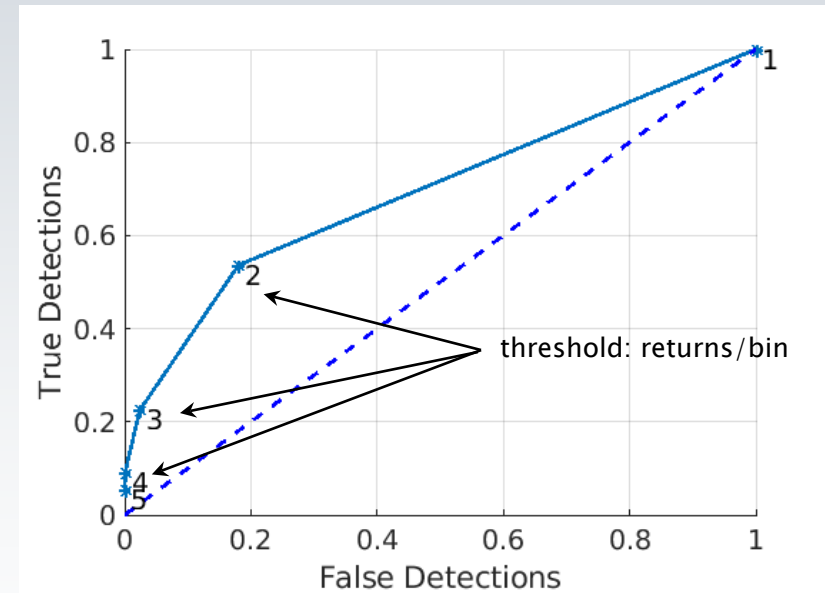
Quantitative Analysis: Receiver Operating Characteristic Curves

Case A: High Signal-to-Noise

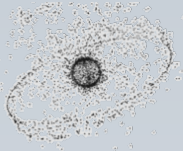


(1 ns x 2s rectangular bins)

Case B: Low Signal-to-Noise



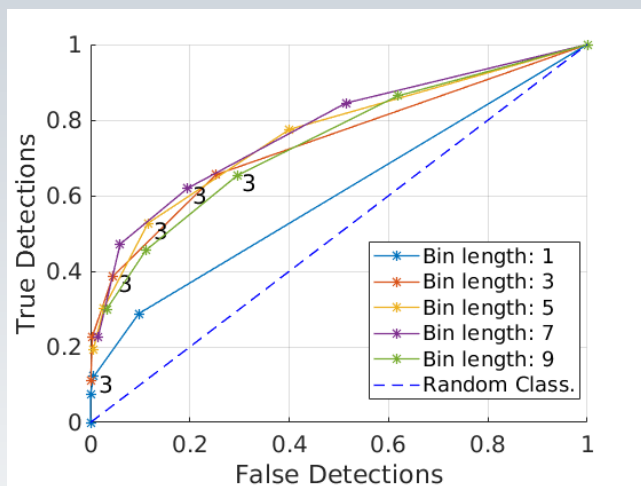
How can we improve the performance?



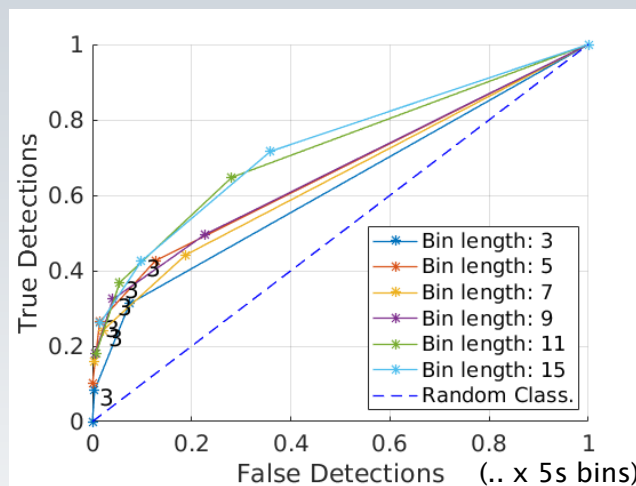
Processing Weak Signals from SDLR

Optimization of Bins

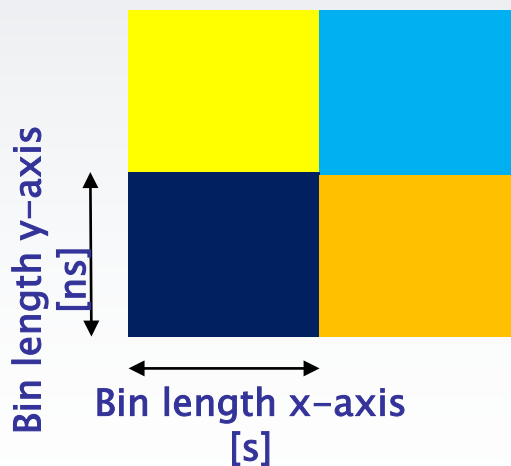
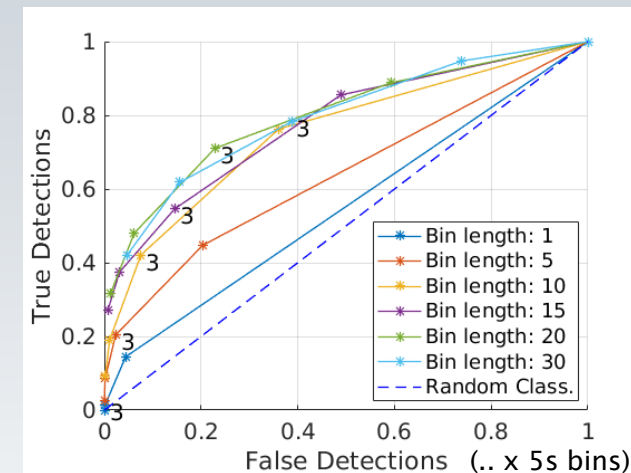
Bin length in x-axis



Bin length in y-axis

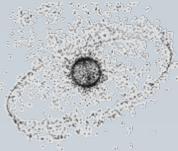


Bin shape: hexagons



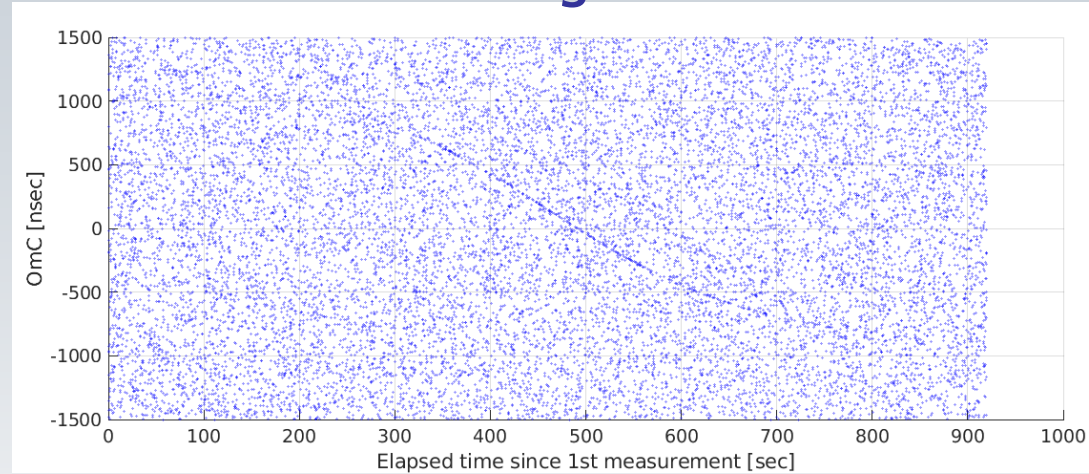
- Reduction of sampling bias → Edge effect
- Low perimeter-to-area ratio → No overlaps or gaps
- Capability to adapt to curve patterns
- Correlations between all possible combinations



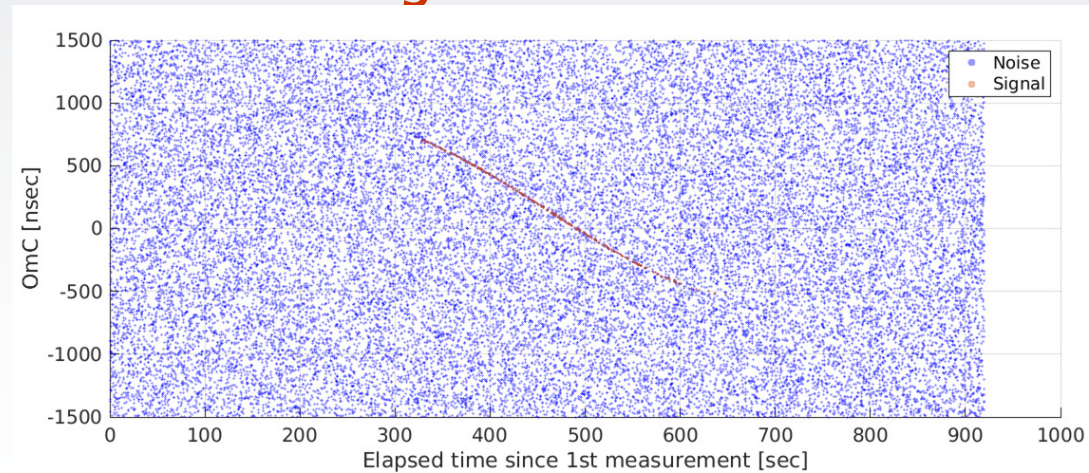


Processing Weak Signals from SDLR

Case B: Low Signal-to-Noise



Filtered signal with new method



Summary

- A space debris laser is currently installed in the Coudé path of the ZIMLAT telescope
 - 40W 100/200 Hz
- A monostatic Coudé configuration poses specific challenges
 - 10^{-5} attenuation for calibration needed
 - laser repetition rate: 6 separate working points for laser
- New algorithms for processing weak signals for SDLR were developed and validated by J. Rodriguez

Thank you for your time!

